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3 In order to obtain the field equations of  $f(T, f)$  gravity to study its cosmological applications, we consider the homogeneous and isotropic Friedmann-Lemaître-Robertson-Walker (FLRW) space time as  $ds^2 = dt^2 + a^2(t)[dx^2 + dy^2 + dz^2]$ , (6) where  $a(t)$  is the scale factor that represents the expansion rate in the spatial directions and the tetrad,  $e_A^m$

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